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AUGUST 2008 NUMBER 768 - Part II

# **Solar-Geophysical Data comprehensive reports**

Data for February 2008

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## **NATIONAL GEOPHYSICAL DATA CENTER**

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Boulder, Colorado

AUGUST 2008 NUMBER 768 - Part II



# Solar-Geophysical Data comprehensive reports

Data for February 2008 and Miscellaneous  
Explanation of Data Reports Issued as Number 515 (Supplement) July 1987

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NATIONAL OCEANIC AND  
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NATIONAL GEOPHYSICAL  
DATA CENTER

BOULDER,  
COLORADO

## DETAILED INDEX OF OBSERVATIONS PUBLISHED IN SOLAR-GEOPHYSICAL DATA

CODE	KIND OF OBSERVATION	DEC	JAN 08	FEB	MAR	APR	MAY	JUN	JUL
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The entry "748A 48" under Oct, for example, means the sunspot drawings for Oct appear in SOLAR-GEOPHYSICAL DATA No 748, Part I, and that they begin on page 48, "A" denotes Part I and "B", Part II. Blanks indicate data not yet received and dashes mark unavailable data.

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HÀ S O L A R F L A R E S  
F E B R U A R Y 2 0 0 8

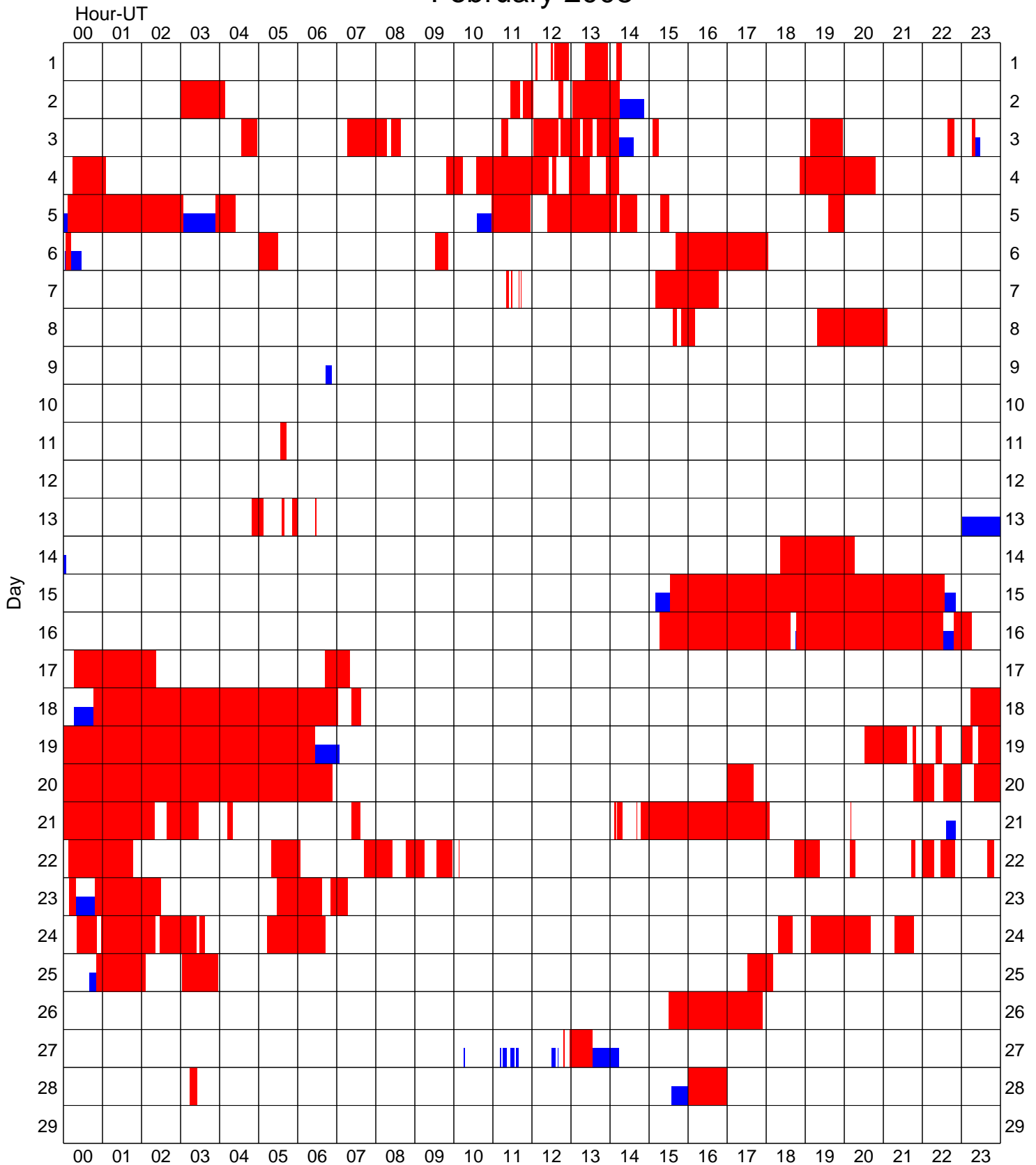
Grp #	Sta	Day	Start (UT)	Max (UT)	End (UT)	Lat	CMD	NOAA/USAF		Dur (Min)	Imp Opt	Xray	Obs See	Type	Area Measurement		Remarks
								Region	Mo Day						Time (UT)	Apparent (10-6 Disk)	
	19		2326		2400	No	Flare	Patrol									
	20		0000		0653	No	Flare	Patrol									
	20		1700		1740	No	Flare	Patrol									
	20		2146		2218	No	Flare	Patrol									
	20		2232		2258	No	Flare	Patrol									
	20		2320		2400	No	Flare	Patrol									
	21		0000		0220	No	Flare	Patrol									
	21		0239		0328	No	Flare	Patrol									
	21		0412		0420	No	Flare	Patrol									
	21		0723		0736	No	Flare	Patrol									
	21		1406		1409	No	Flare	Patrol									
	21		1411		1419	No	Flare	Patrol									
	21		1441		1442	No	Flare	Patrol									
	21		1447		1805	No	Flare	Patrol									
	21		2010		2011	No	Flare	Patrol									
	22		0008		0147	No	Flare	Patrol									
	22		0519		0604	No	Flare	Patrol									
	22		0742		0825	No	Flare	Patrol									
	22		0846		0914	No	Flare	Patrol									
	22		0934		0958	No	Flare	Patrol									
	22		1008		1009	No	Flare	Patrol									
	22		1843		1922	No	Flare	Patrol									
	22		2009		2017	No	Flare	Patrol									
	22		2143		2149	No	Flare	Patrol									
	22		2200		2218	No	Flare	Patrol									
	22		2228		2250	No	Flare	Patrol									
	22		2340		2350	No	Flare	Patrol									
	23		0009		0019	No	Flare	Patrol									
	23		0049		0230	No	Flare	Patrol									
	23		0528		0637	No	Flare	Patrol									
	23		0651		0717	No	Flare	Patrol									
	24		0021		0051	No	Flare	Patrol									
	24		0058		0221	No	Flare	Patrol									
	24		0228		0324	No	Flare	Patrol									
	24		0329		0337	No	Flare	Patrol									
	24		0513		0643	No	Flare	Patrol									
	24		1818		1840	No	Flare	Patrol									
	24		1909		2041	No	Flare	Patrol									
	24		2118		2147	No	Flare	Patrol									
	25		0050		0206	No	Flare	Patrol									
	25		0303		0358	No	Flare	Patrol									
	25		1731		1810	No	Flare	Patrol									
	26		1530		1754	No	Flare	Patrol									
	27		1248		1250	No	Flare	Patrol									
	27		1258		1333	No	Flare	Patrol									
	28		0314		0325	No	Flare	Patrol									
	28		1601		1700	No	Flare	Patrol									

"Remarks"

- |   |  |
|---|--|
| A = Eruptive prominence whose base is less than 90 degrees from central meridian. | O = Observations have been made in the H and K lines of Ca II.   |
| B = Probably the end of a more important flare.                                   | P = Flare shows Helium D3 in emission.   |
| C = Invisible 10 minutes before.  | Q = Flare shows Balmer continuum in emission.  |
| D = Brilliant point.  | R = Marked asymmetry in H-alpha line suggests ejection of high-velocity material.  |
| E = Two or more brilliant points.   | S = Brightness follows disappearance of filament in same position.   |
| F = Several eruptive centers.   | T = Region active all day.   |
| G = No visible spots in the neighborhood.   | U = Two bright branches, parallel or converging.   |
| H = Flare accompanied by high-speed dark filament.                                | V = Occurrence of an explosive phase; important, expansion within roughly 1 minute that often includes a significant intensity increase. |
| I = Active region very extended.  | W = Great increase in area after time of maximum intensity.  |
| J = Distinct variations of plage intensity before or after the flare.             | X = Unusually wide H-alpha line.   |
| K = Several intensity maxima.   | Y = System of loop-type prominences.   |
| L = Existing filaments show signs of sudden activity.                             | Z = Major sunspot umbra covered by flare.  |
| M = White-light flare.  |  |
| N = Continuous spectrum shows effects of polarization.                            |  |

Observation Type: C=Cinematographic, E=Electronic, P=Photographic, V=Visual

# Intervals of No Flare Patrol Observation for Preceding Solar Flare Table February 2008

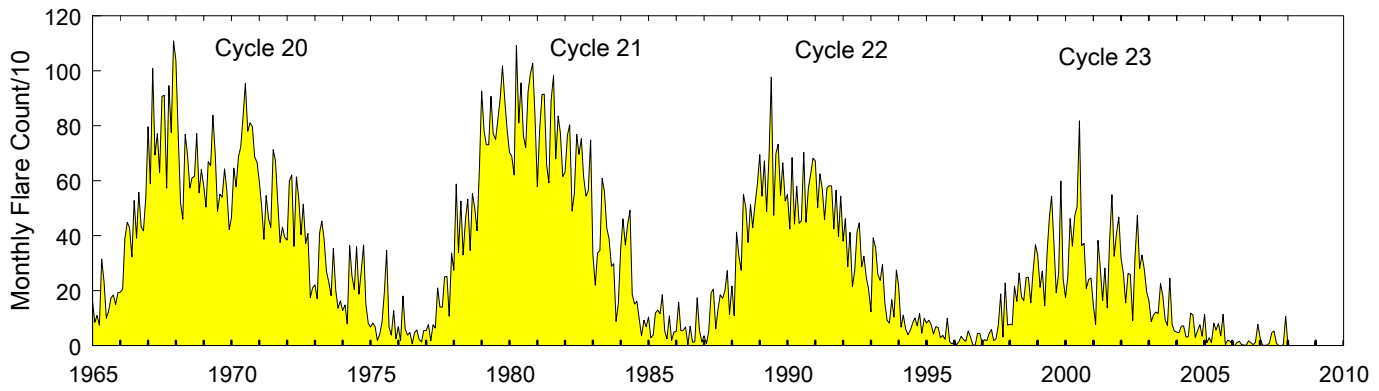


■ Times of no flare patrol of any kind.  
■ Times of no cinematographic flare patrol.

Stations participating: Holloman, Learmonth, SanVito, Kanzelhoehe.

# Monthly Counts of Grouped Solar Flares

## Jan 1965 - Feb 2008



Year	Jan	Feb	Mar	Apr	May	Jun	Jul	Aug	Sep	Oct	Nov	Dec	Total
1965	158	85	110	74	315	231	99	127	173	184	150	193	1899
1966	194	205	390	449	429	323	528	391	558	432	417	543	4859
1967	796	589	1009	694	771	629	907	911	573	946	775	1109	9709
1968	1037	773	519	460	768	697	573	611	616	772	556	640	8022
1969	581	504	669	655	839	694	489	551	540	643	566	422	7153
1970	466	646	578	688	722	836	954	780	811	797	687	667	8632
1971	598	505	387	546	461	430	713	673	518	375	431	394	6031
1972	384	599	621	361	614	541	404	515	371	408	175	210	5203
1973	221	171	410	453	388	270	232	182	353	201	136	163	3180
1974	127	148	79	364	255	204	360	187	270	366	153	81	2594
1975	68	82	69	19	42	85	196	346	68	38	127	25	1165
1976	69	18	180	60	38	48	6	47	57	23	13	55	614
1977	54	77	18	76	64	210	140	140	250	252	107	336	1724
1978	274	588	338	526	330	460	533	346	554	499	418	648	5514
1979	926	781	731	731	907	772	750	821	901	1018	888	786	10012
1980	703	689	621	1092	811	956	763	720	924	988	1027	838	10132
1981	578	782	914	915	658	592	893	982	680	836	773	615	9218
1982	631	766	803	490	553	769	696	753	615	544	564	748	7932
1983	332	220	337	346	609	561	427	389	289	298	88	152	4048
1984	353	461	366	440	492	185	151	161	95	36	92	69	2901
1985	104	29	38	119	129	116	185	53	25	108	19	50	975
1986	51	158	54	56	68	3	71	12	14	174	56	13	730
1987	36	7	52	192	205	61	132	185	172	198	273	114	1627
1988	217	109	413	328	274	551	502	375	513	429	518	587	4816
1989	695	544	672	488	691	977	474	699	733	547	665	526	7711
1990	550	424	684	442	580	445	454	703	449	574	623	682	6610
1991	672	503	625	570	458	574	582	581	425	565	396	544	6495
1992	380	462	287	412	214	271	413	447	287	325	248	206	3952
1993	123	392	357	262	237	296	154	92	82	167	104	275	2541
1994	217	67	111	60	40	56	81	101	72	117	45	99	1066
1995	82	95	77	42	69	66	29	37	23	99	14	6	639
1996	14	3	15	34	21	16	54	31	3	0	44	45	280
1997	8	22	18	43	59	18	26	75	188	31	228	74	790
1998	78	76	216	161	264	177	164	248	249	155	268	367	2423
1999	330	212	271	145	330	466	544	368	192	264	598	243	3963
2000	175	248	462	362	473	505	818	364	372	208	241	246	4474
2001	147	77	383	284	164	282	137	376	549	325	405	468	3597
2002	318	261	155	263	259	91	318	474	280	329	279	196	3223
2003	164	87	112	122	117	226	181	94	73	245	78	53	1552
2004	49	47	71	72	32	33	118	112	30	54	76	34	728
2005	114	10	28	11	82	56	81	35	114	4	20	16	571
2006	4	0	11	16	4	2	1	17	11	3	12	78	159
2007	29	2	1	2	9	47	53	9	0	0	2	107	261
2008	2	0											2

The term 'grouped' means observations of the same event by different sites were lumped together and counted as one.



8  
Feb 08

S O L A R R A D I O E M I S S I O N  
Outstanding Occurrences  
FEBRUARY 2008

Day	Freq Sta	Type	Start (UT)	Time of Maximum (UT)	Duration (Min)	Flux Density		Int	Remarks
						Peak (10 -22 W/m <sup>2</sup> Hz)	Mean		
02	245 LEAR	8 S	2309.0	2309.0	U	160.0			QL=4 ST=2 TYP=3
07	33 UPIC	3 S	1003.0	1003.2	0.5U				
08	245 LEAR	8 S	0927.0	0927.0	U	140.0			QL=4 ST=2 TYP=3
27	33 UPIC	3 S	1258.5	1258.7	0.5U				

Reports are received routinely from the following observatories:

CUBA = Havana	LEAR = Learmonth	SGMR = Sagamore Hill
GORK = Gorky	PEKG = Peking	SVTO = San Vito
HIRA = Hiraiso	PALE = Palehua	TORN = Torun
IZMI = IZMIRAN	PENT = Penticton	UPIC = Upice

Explanation of Type Code:

1 Simple 1	7 Minor +	24 Rise	30 Post Burst Increase A	43 Onset of Noise Storm
2 Simple 1F	8 Spike	25 Rise A	31 Post Burst Decrease	44 Noise Storm in Progress
3 Simple 2	20 Simple 3	26 Fall	33 Absorption	45 Complex
4 Simple 2F	21 Simple 3A	27 Rise and Fall	40 Fluctuation	46 Complex F
5 Simple	22 Simple 3F	28 Precursor	41 Group of Bursts	47 Great Burst
6 Minor	23 Simple 3AF	29 Post Burst Increase	42 Series of Bursts	48 Major
1A Simple 1A	4A Simple 2AF	24PF Post Rise F	27F Rise and Fall F	
3A Simple 2A	40 Rise Only	16A Fall A	27AF Rise and Fall AF	
21A Simple 3A GRF	40F Rise Only F	260 Fall Only	31A Post Burst Decrease A	
2A Simple 1AF	4P Post Rise	26F Fall F	32A Absorption A	

RSTN Site Information: Beginning in April 1986, the RSTN sites LEAR, PALE, SGMR, and SVTO fixed frequency solar radio data are periodically adjusted to several world standard stations. These world standard stations include: Kislovodsk, USSR 15,500 MHz; Penticton, Canada 2800 MHz; and Hiraiso, Japan 500 and 200 MHz.

## GOES Space Environment Monitor

The Geostationary Operational Environmental Satellites (GOES-1, GOES-2, etc.) all carry on board the Space Environment Monitor (SEM) instrument subsystem. The SEM has provided magnetometer, energetic particle, and soft X-ray data continuously since July 1974.

Geosynchronous satellites have an unobstructed view of the sun for all but the few dozen hours per year when the Earth eclipses the sun. You can identify these intervals as gaps in the X-ray data near satellite local midnight in March-April, and September-October.

The source data for these plots are averages; as a result, the peak values may appear smaller than they actually were.

The volume of these data makes it impossible to issue a guarantee as to the quality of each and every data point. Users should be suspicious of 'spikes' in the data and attempt to correlate them with other sources before assuming that they represent the space environment. The time of these observations has not been corrected for the downlink and preprocessing delay which is within 1 - 5 seconds.

## X-ray Sensor (XRS)

Ion chamber detectors provide whole-Sun X-ray fluxes for the 0.5-to-3 (0.5-to-4 prior to GOES-8) and 1-to-8 Å wavelength bands. The X-ray sensors may experience significant bremsstrahlung contamination caused by energetic particles in the outer radiation belts and depends on satellite local time, time of year, and the local particle pitch-angle distribution. The X-ray sensors are also sensitive to background contamination due to energetic electrons that either deposit their energy directly in the telescope or strike the external structure and produce bremsstrahlung X-rays inside the ion chambers.

## Solar Flares

A solar flare is a short-lived sudden increase in the intensity of solar radiation that originates at or near sunspots. Ground based optical observations are best made at the emission line of Hydrogen known as H-alpha (6563 Å). Thanks to satellite born instrumentation, flares can now be observed at X-ray wavelengths (0.1-10.0 Å). Flares are characterized by a rise time of the order of minutes and decay of the order of tens of minutes. The total energy expended in a typical flare is about  $10^{30}$  ergs; the magnetic field is extraordinarily high, reaching values of 100 to 10,000 gauss. Optical flares are usually accompanied by radio and X-ray bursts, and occasionally by high-energy particle emissions. The National Geophysical Data Center (NGDC) holds archives for about 80 stations, covering the period 1937 to the present. Currently about 5 stations send their data to NGDC Boulder on a routine monthly basis. The reports are processed and published in the monthly report "Solar-Geophysical Data" and in a different format in the IAU "Quarterly Bulletin on Solar Activity."

## X-ray Flare Reports

An X-ray flare officially begins when four consecutive 1-minute X-ray values meet the following conditions:

1. All four values exceed the B threshold ( $10^{-7}$  Watts/m<sup>2</sup>)
  2. All four values are strictly increasing.
  3. The last value is 1.4 times greater than the first value.
- The X-ray flare classification is based on power per area (Watts/m<sup>2</sup>) at the time of maximum. A letter value (B, C, M or X) represents the decadal flux value as shown in the data key below. The letter value is followed by a number which tells us the specific intensity of the flare. For example, a flare classified as M3.5 has a peak intensity of  $3.5 \times 10^{-5}$  Watts/m<sup>2</sup>. The event ends when the flux reading drops below half the sum of the maximum and starting flux. Flare start times are marked on the plot with a " | ".

## H-alpha Flare Reports

The basic reports sent monthly from the observatories consist of data for each flare or subflare observation by photographic or visual patrol. The data files give as many of the following measurements as possible: time of beginning; time of maximum brightness; time of any prominent secondary maxima; time of end (all times in UT); area at time of maximum brightness (square degrees of solar disk correct for foreshortening); importance class of flare (IAU 1964 report, updated in 1975); heliographic coordinates of center of gravity of flare at maximum brightness. Flare start times are marked on the plot with a " | ".

## Data Key

- XL** 1 - 8 Å X-rays (Watts/m<sup>2</sup>)
- XS** 0.5 - 3 Å X-rays, or 0.5 - 4 Å prior to GOES-8 (Watts/m<sup>2</sup>)
- | X-ray and H-alpha flare report start times.
- X** X-ray flux  $> 10^{-4}$  Watts/m<sup>2</sup>
- M** X-ray flux  $10^{-5} - 10^{-4}$  Watts/m<sup>2</sup>
- C** X-ray flux  $10^{-6} - 10^{-5}$  Watts/m<sup>2</sup>
- B** X-ray flux  $10^{-7} - 10^{-6}$  Watts/m<sup>2</sup>

## Contact Information

These plots were generated by NOAA's National Geophysical Data Center:  
<http://www.ngdc.noaa.gov>

GOES SEM Data: <http://www.ngdc.noaa.gov/stp/GOES/goes.html>

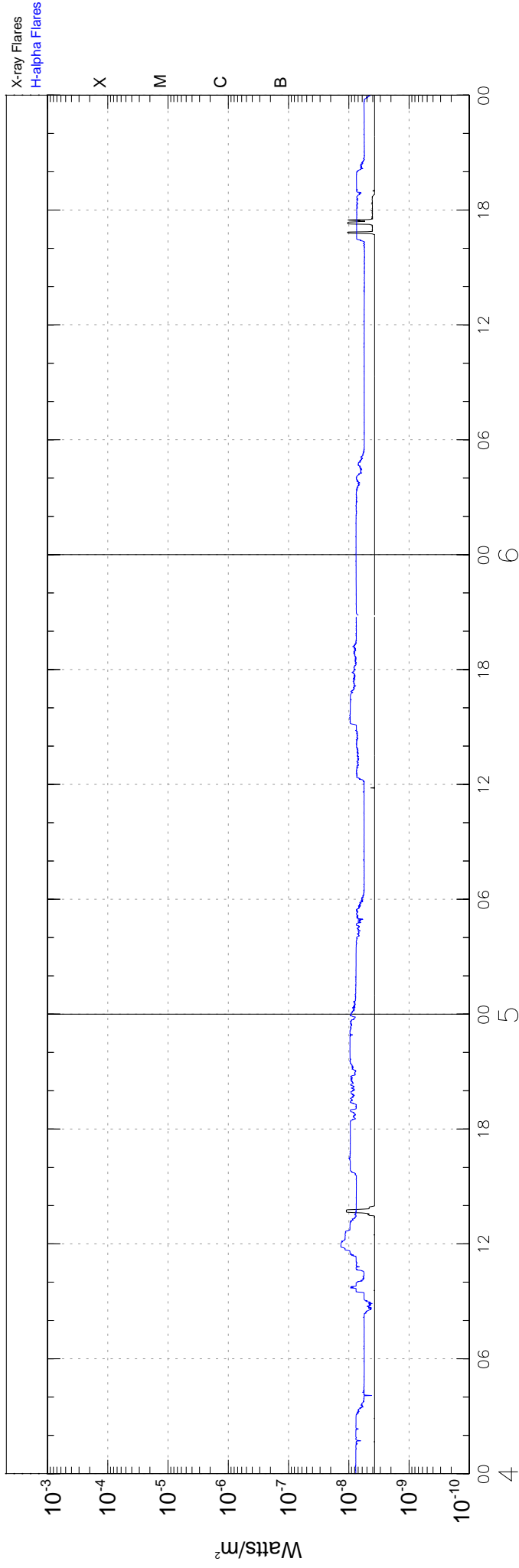
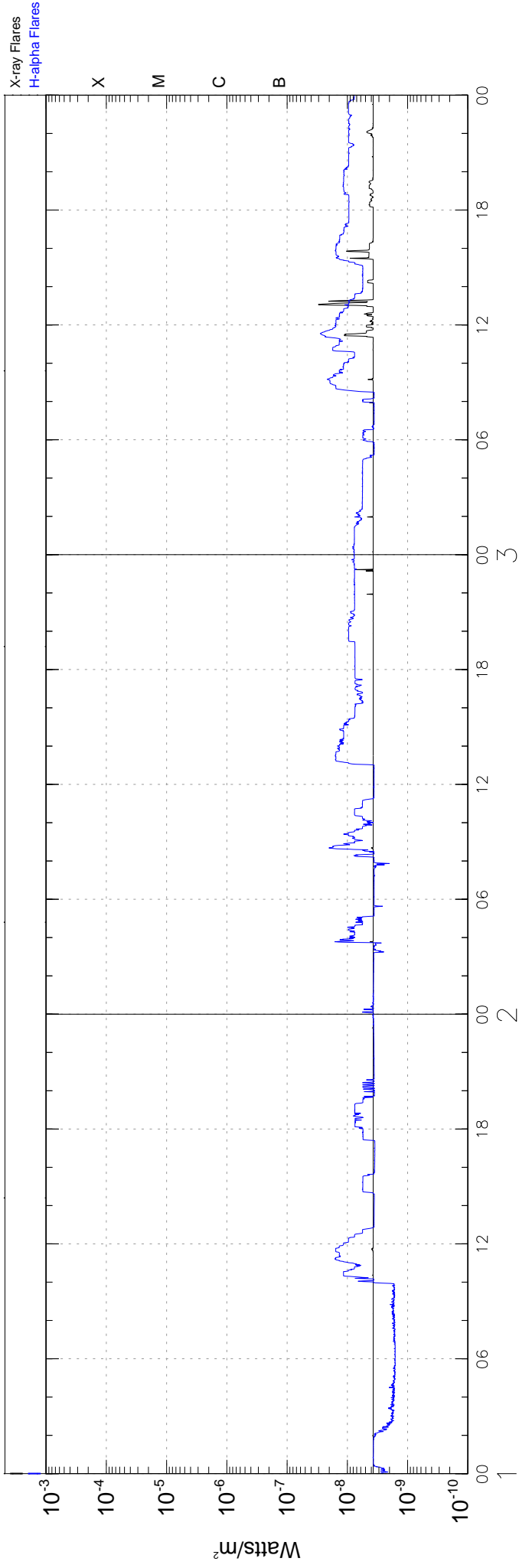
Flare Reports: <http://www.ngdc.noaa.gov/stp/SOLAR/ftp/solarflares.html>

GOES SEM data are processed by NOAA's Space Environment Center:  
<http://www.swpc.noaa.gov/>

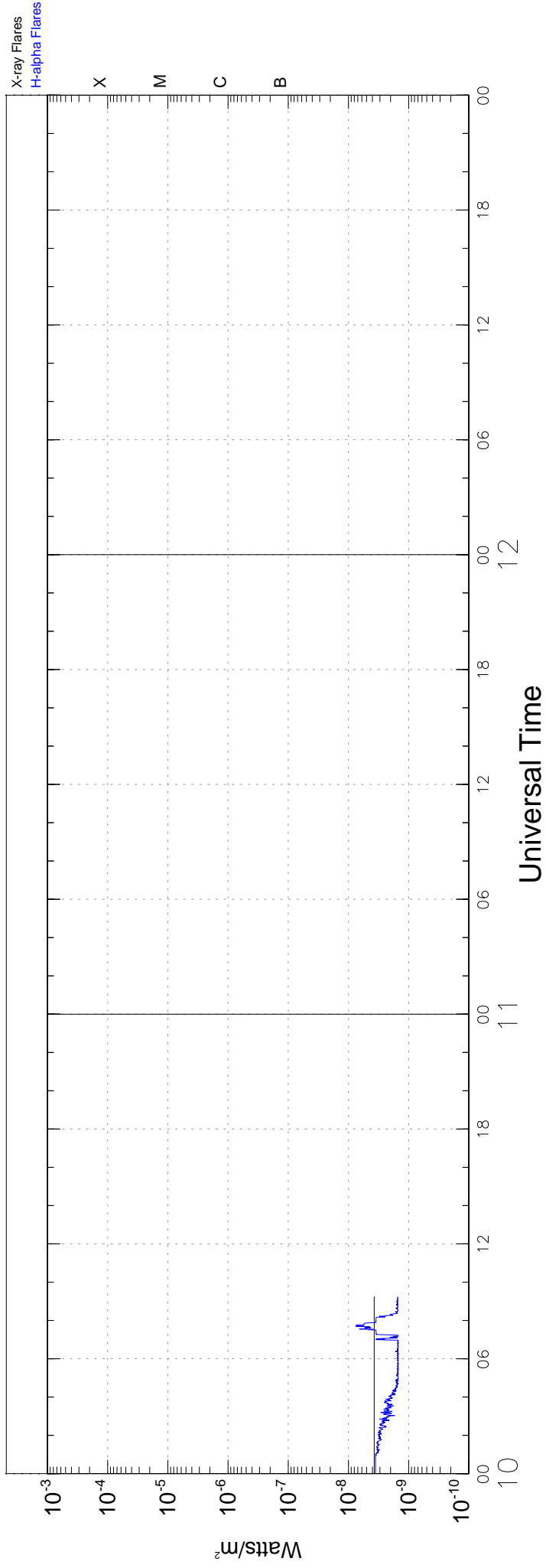
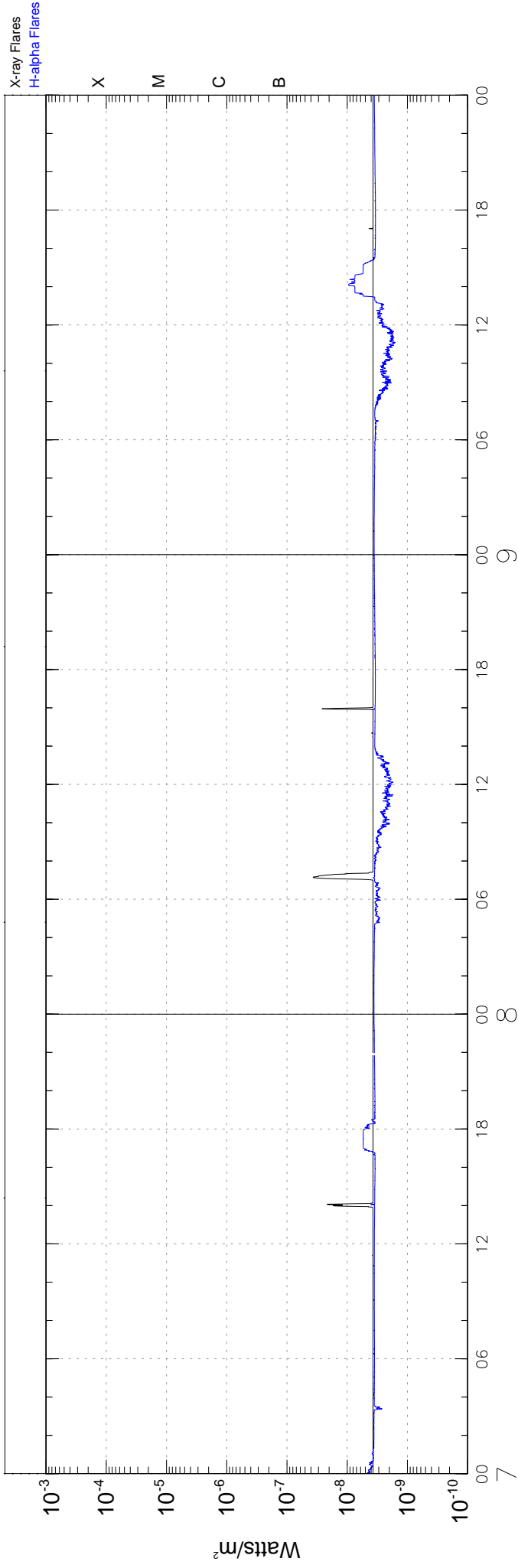
Address comments to: [Daniel.C.Wilkinson@noaa.gov](mailto:Daniel.C.Wilkinson@noaa.gov) or [Edward.H.Erwin@noaa.gov](mailto:Edward.H.Erwin@noaa.gov)



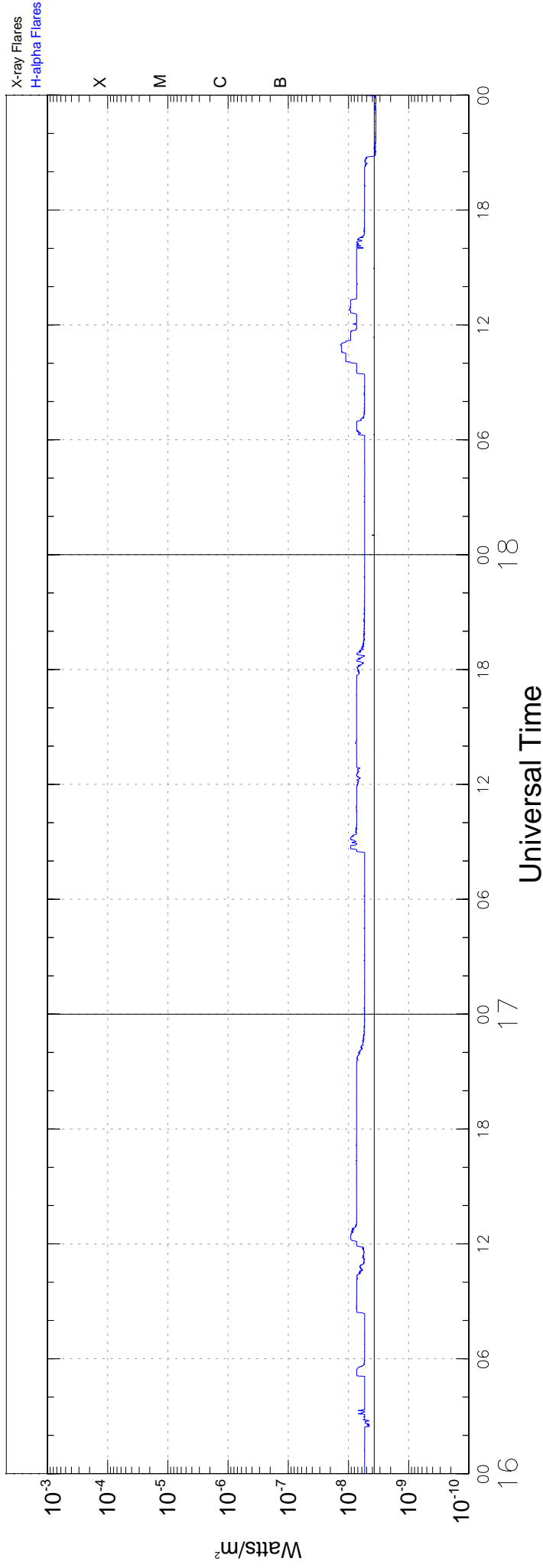
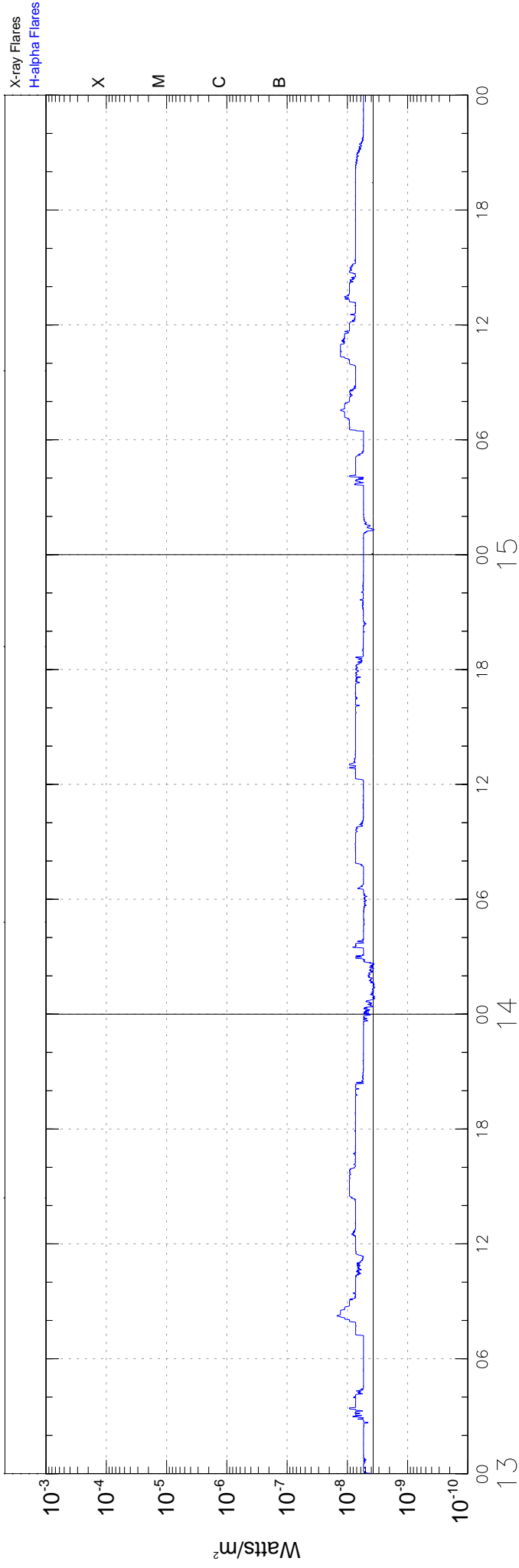
# GOES-11 Solar X-Rays (1-Minute Averages) February 2008



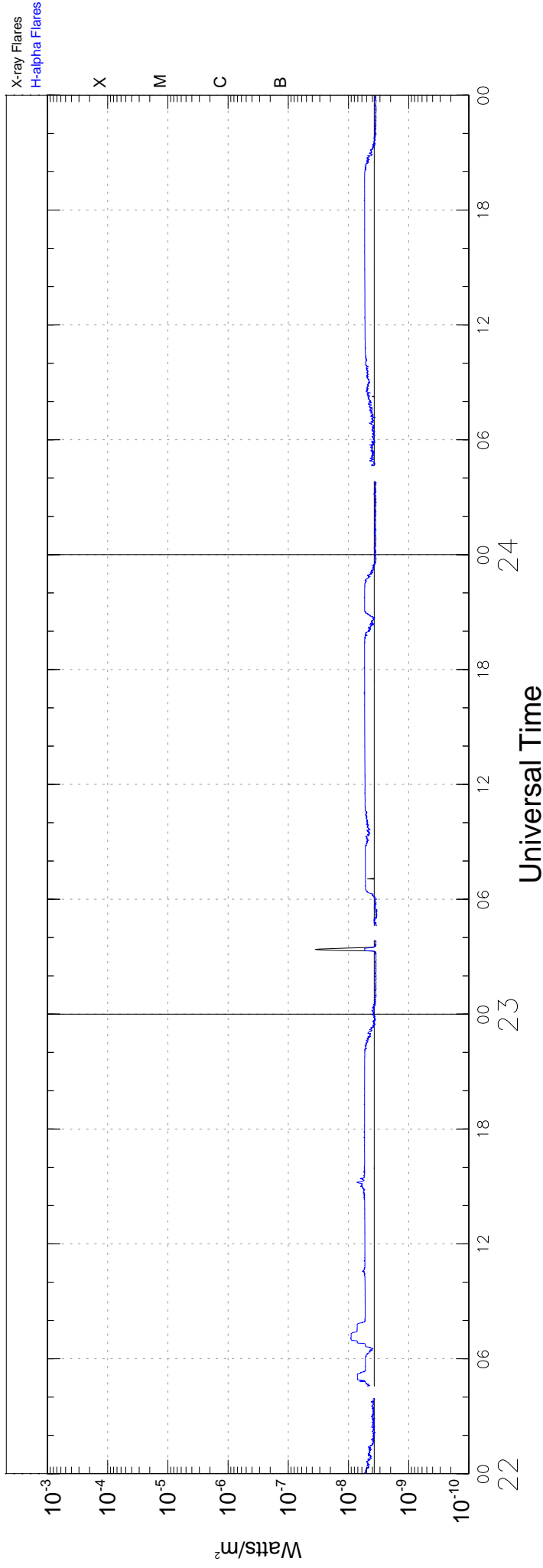
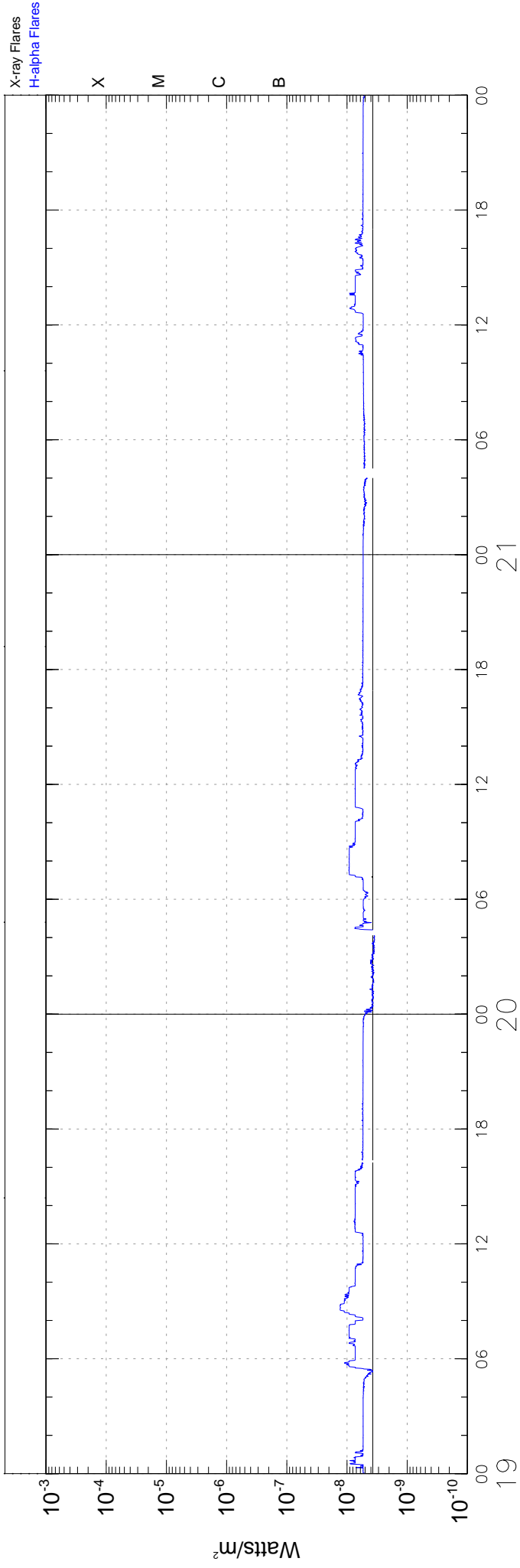
# GOES-11 Solar X-Rays (1-Minute Averages) February 2008



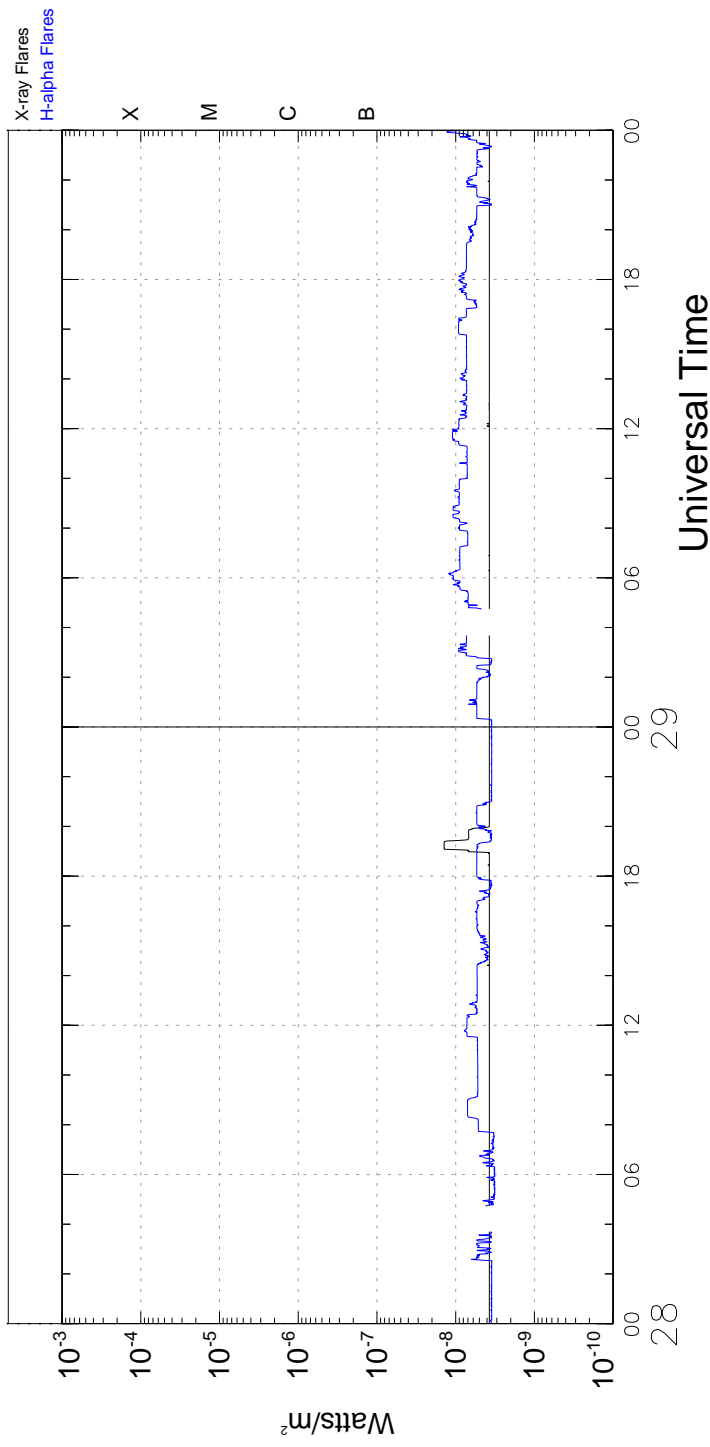
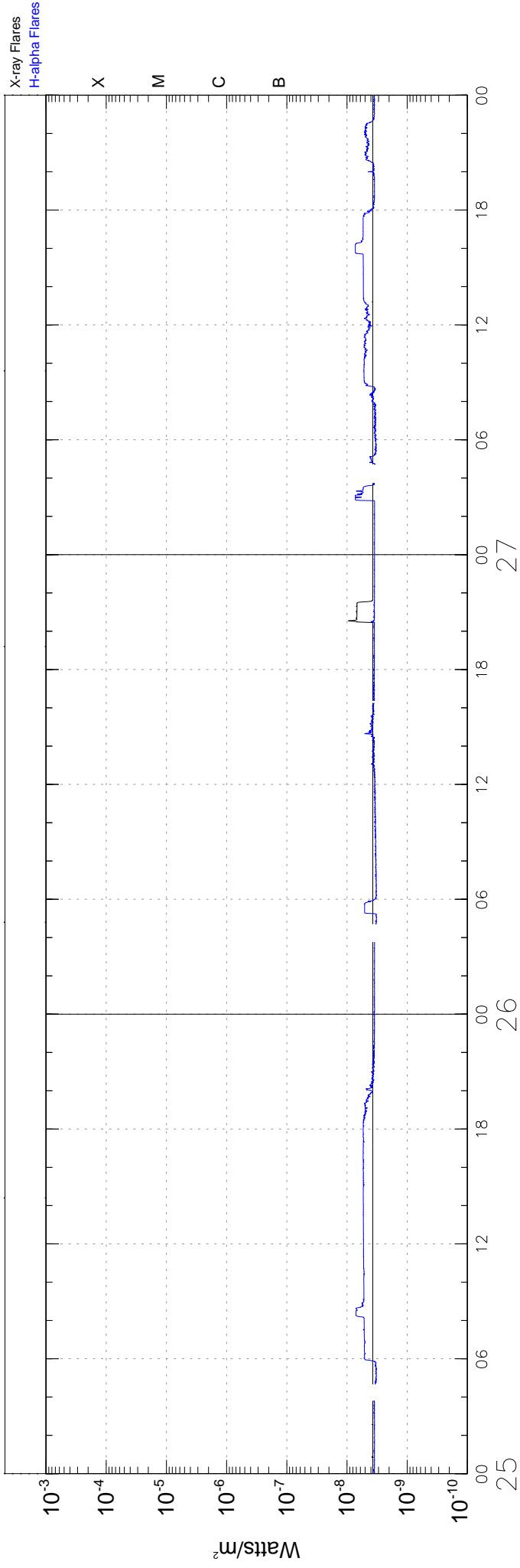
# GOES-10 Solar X-Rays (1-Minute Averages) February 2008



# GOES-10 Solar X-Rays (1-Minute Averages) February 2008



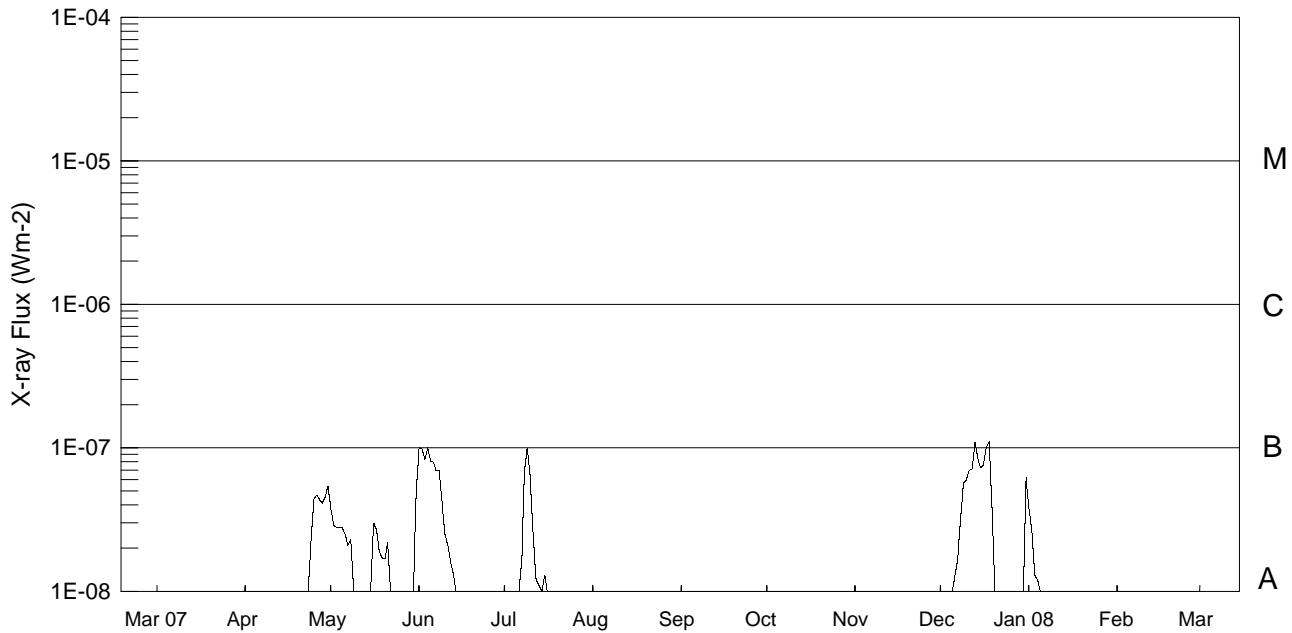
# GOES-10 Solar X-Rays (1-Minute Averages) February 2008



# Preliminary GOES Satellite Daily X-Ray Background

## Mar 2007 - Feb 2008

15  
Feb 08



Day	Mar 07	Apr	May	Jun	Jul	Aug	Sep	Oct	Nov	Dec	Jan 08	Feb
1	<A1.0	<A1.0	A3.8	B1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	A3.8	<A1.0
2	<A1.0	<A1.0	A2.9	A9.9	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	A2.7	<A1.0
3	<A1.0	<A1.0	A2.8	A8.3	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	A1.3	<A1.0
4	<A1.0	<A1.0	A2.8	B1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	A1.2	<A1.0
5	<A1.0	<A1.0	A2.8	A8.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	A1.0	<A1.0
6	<A1.0	<A1.0	A2.5	A8.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	A1.3	<A1.0	<A1.0
7	<A1.0	<A1.0	A2.1	A7.0	A1.7	<A1.0	<A1.0	<A1.0	<A1.0	A1.7	<A1.0	<A1.0
8	<A1.0	<A1.0	A2.3	A7.0	A6.9	<A1.0	<A1.0	<A1.0	<A1.0	A3.2	<A1.0	<A1.0
9	<A1.0	<A1.0	<A1.0	A4.2	B1.0	<A1.0	<A1.0	<A1.0	<A1.0	A5.7	<A1.0	<A1.0
10	<A1.0	<A1.0	<A1.0	A2.5	A6.3	<A1.0	<A1.0	<A1.0	<A1.0	A6.0	<A1.0	<A1.0
11	<A1.0	<A1.0	<A1.0	A2.1	A2.4	<A1.0	<A1.0	<A1.0	<A1.0	A7.0	<A1.0	<A1.0
12	<A1.0	<A1.0	<A1.0	A1.6	A1.2	<A1.0	<A1.0	<A1.0	<A1.0	A7.1	<A1.0	<A1.0
13	<A1.0	<A1.0	<A1.0	A1.3	A1.1	<A1.0	<A1.0	<A1.0	<A1.0	B1.1	<A1.0	<A1.0
14	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	A8.4	<A1.0	<A1.0
15	<A1.0	<A1.0	<A1.0	<A1.0	A1.3	<A1.0	<A1.0	<A1.0	<A1.0	A7.3	<A1.0	<A1.0
16	<A1.0	<A1.0	A3.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	A7.5	<A1.0	<A1.0
17	<A1.0	<A1.0	A2.7	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	B1.0	<A1.0	<A1.0
18	<A1.0	<A1.0	A1.9	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	B1.1	<A1.0	<A1.0
19	<A1.0	<A1.0	A1.7	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	A4.0	<A1.0	<A1.0
20	<A1.0	<A1.0	A1.7	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0
21	<A1.0	<A1.0	A2.2	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0
22	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0
23	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0
24	<A1.0	A2.2	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0
25	<A1.0	A4.4	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0
26	<A1.0	A4.7	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0
27	<A1.0	A4.4	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0
28	<A1.0	A4.1	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0
29	<A1.0	A4.5	<A1.0	A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0
30	<A1.0	A5.4	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0	<A1.0
31	<A1.0		A5.1		<A1.0	<A1.0		<A1.0		A6.2	<A1.0	

Levels below B1.0 are unreliable.

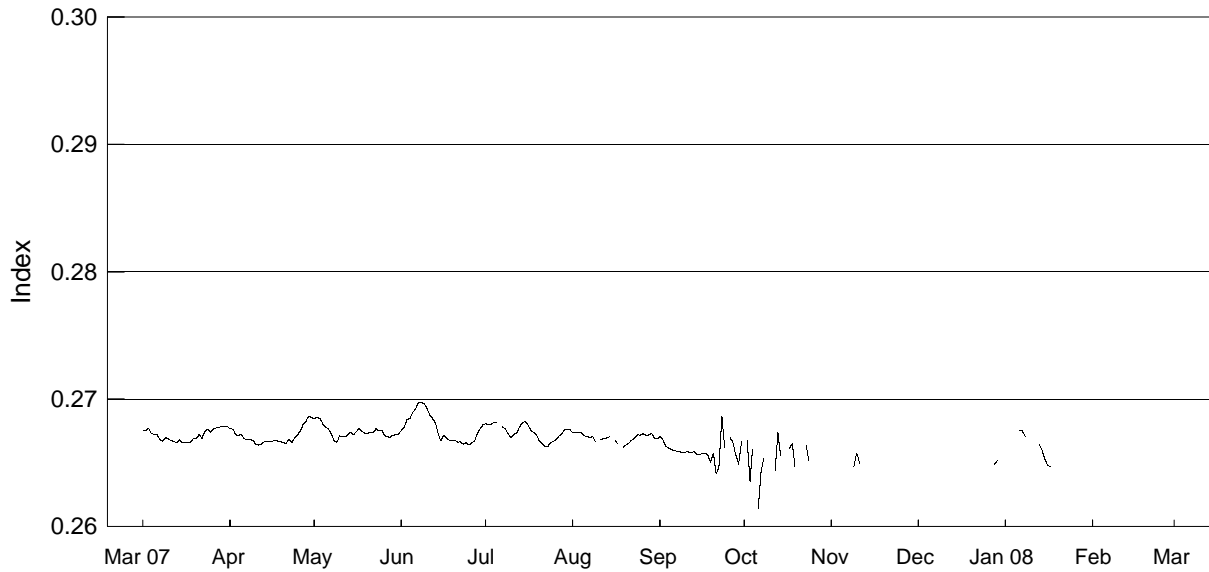




# NOAA Solar Ultraviolet (UV) MgII Core-to-Wing Index

## Mar 2007 - Feb 2008

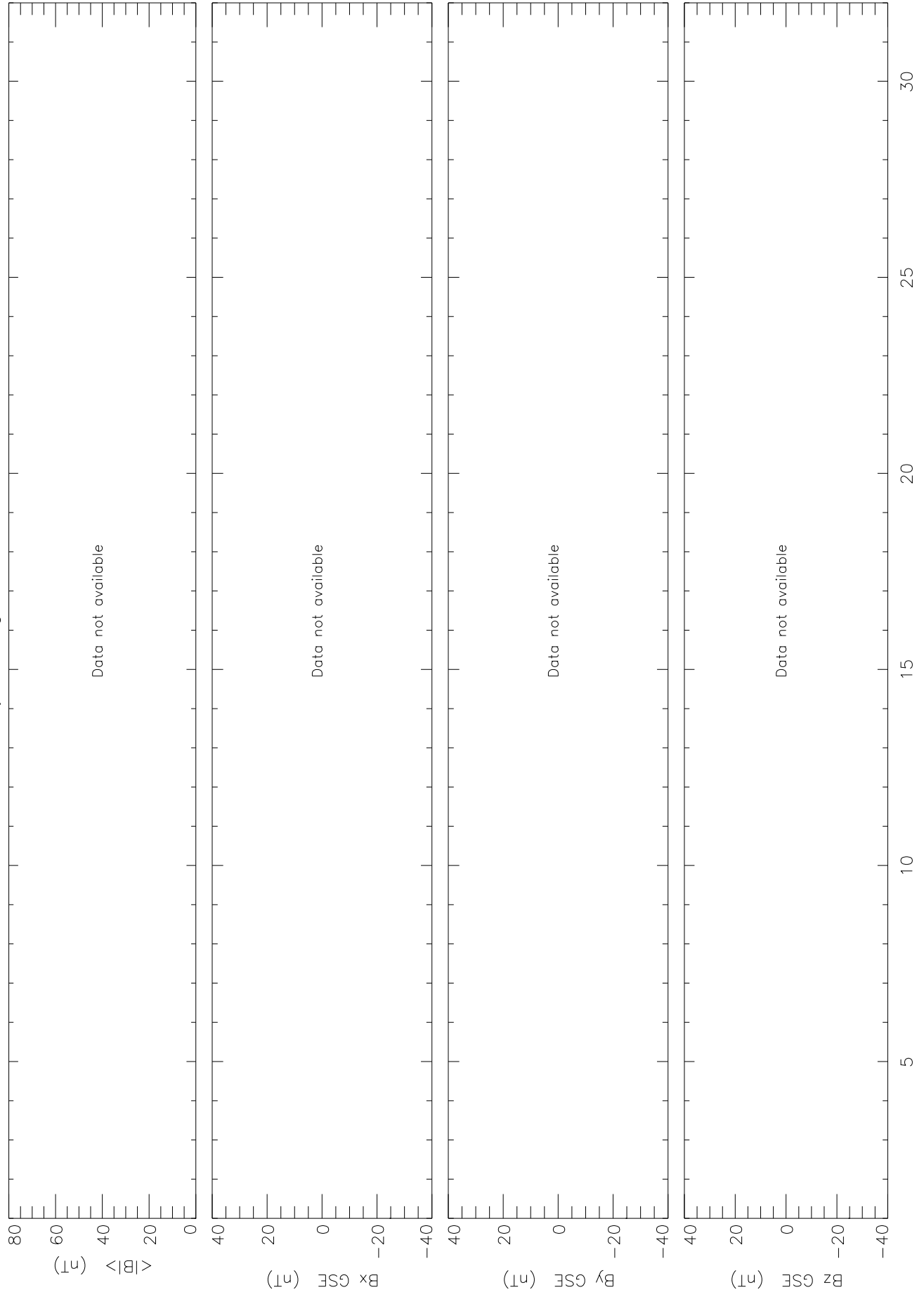
### Version 9.1



Day	Mar 07	Apr	May	Jun	Jul	Aug	Sep	Oct	Nov	Dec	Jan 08	Feb
1	0.2675	0.2677	0.2685	0.2676	0.2681	0.2674	0.2671	---	---	---	---	---
2	0.2675	0.2676	0.2686	0.2678	0.2680	0.2674	0.2669	0.2668	---	---	0.2658	---
3	0.2677	0.2672	0.2685	0.2684	0.2681	0.2674	0.2663	0.2635	---	---	---	---
4	0.2673	0.2671	0.2680	0.2685	0.2682	0.2674	0.2662	0.2661	---	---	---	---
5	0.2673	0.2673	0.2679	0.2690	0.2682	0.2672	0.2661	---	---	---	---	---
6	0.2672	0.2669	0.2677	0.2692	---	0.2671	0.2660	0.2615	---	---	0.2675	---
7	0.2669	0.2669	0.2673	0.2697	0.2678	0.2670	0.2659	0.2645	0.2644	---	0.2675	---
8	0.2667	0.2668	0.2667	0.2698	0.2677	0.2671	0.2659	0.2654	---	---	0.2671	---
9	0.2670	0.2668	0.2666	0.2697	0.2673	0.2667	0.2658	---	0.2647	---	---	---
10	0.2669	0.2665	0.2671	0.2693	0.2670	---	0.2659	---	0.2658	---	---	---
11	0.2668	0.2664	0.2671	0.2688	0.2673	0.2668	0.2659	---	0.2650	---	0.2527	---
12	0.2667	0.2665	0.2671	0.2685	0.2674	0.2669	0.2658	0.2644	---	---	---	---
13	0.2666	0.2666	0.2672	0.2682	0.2678	0.2670	0.2659	0.2674	---	---	0.2664	---
14	0.2668	0.2667	0.2674	0.2674	0.2681	0.2671	0.2657	0.2655	---	---	0.2659	---
15	0.2666	0.2667	0.2672	0.2667	0.2683	---	0.2656	---	---	---	0.2653	---
16	0.2666	0.2667	0.2675	0.2672	0.2680	0.2668	0.2657	---	---	---	0.2648	---
17	0.2666	0.2668	0.2677	0.2670	0.2676	0.2665	0.2657	0.2661	---	---	0.2647	---
18	0.2666	0.2667	0.2675	0.2668	0.2674	---	0.2656	0.2665	---	---	---	---
19	0.2669	0.2667	0.2673	0.2668	0.2672	0.2662	0.2650	0.2647	---	---	---	---
20	0.2669	0.2666	0.2673	0.2668	0.2667	0.2664	0.2657	---	---	---	---	---
21	0.2672	0.2666	0.2674	0.2666	0.2665	0.2666	0.2641	---	---	---	---	---
22	0.2669	0.2668	0.2674	0.2667	0.2663	0.2668	0.2647	---	---	---	---	---
23	0.2675	0.2666	0.2677	0.2665	0.2663	0.2670	0.2687	0.2664	---	---	---	---
24	0.2676	0.2670	0.2675	0.2666	0.2666	0.2672	0.2662	0.2652	---	0.2665	---	---
25	0.2674	0.2672	0.2676	0.2664	0.2667	0.2672	---	---	---	---	---	---
26	0.2677	0.2675	0.2672	0.2666	0.2668	0.2673	0.2670	---	---	---	---	---
27	0.2677	0.2680	0.2671	0.2667	0.2671	0.2672	0.2666	---	---	---	---	---
28	0.2678	0.2683	0.2670	0.2673	0.2673	0.2672	0.2656	---	---	0.2649	---	---
29	0.2678	0.2687	0.2672	0.2677	0.2676	0.2673	0.2649	0.2658	---	0.2652	---	---
30	0.2679	0.2686	0.2672	0.2680	0.2676	0.2669	0.2667	---	---	---	---	---
31	0.2679	---	0.2673	---	0.2676	0.2669	---	---	---	---	---	---
Mean	0.2672	0.2671	0.2674	0.2677	0.2674	0.2670	0.2660	0.2653	0.2650	0.2655	0.2661	---

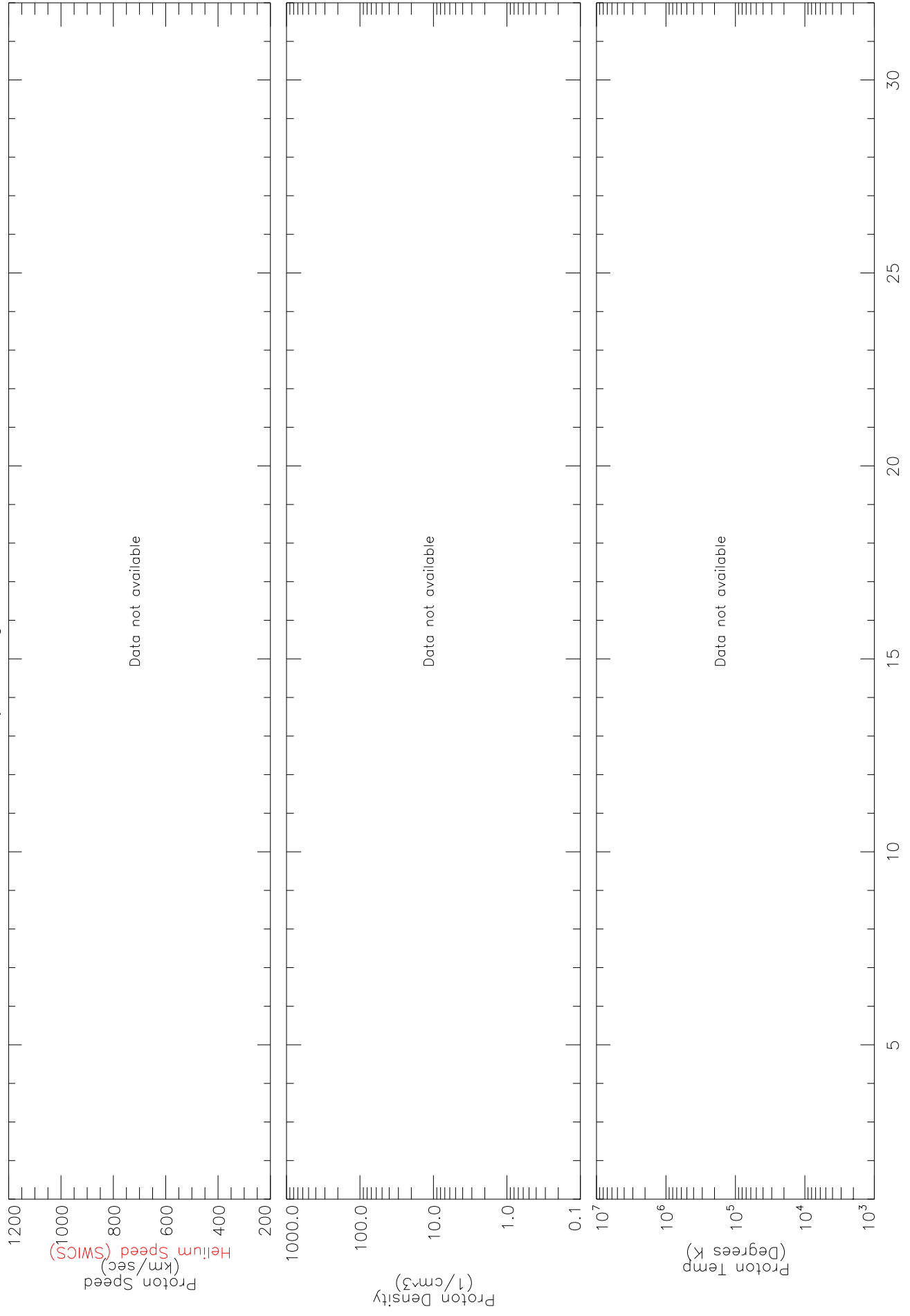
Data at: <http://www.swpc.noaa.gov/ftpmenu/sbuw.html>

ACE LEVEL2 DATA  
Interplanetary Magnetic Field  
Hourly Averages for FEBRUARY 2008, from MAG



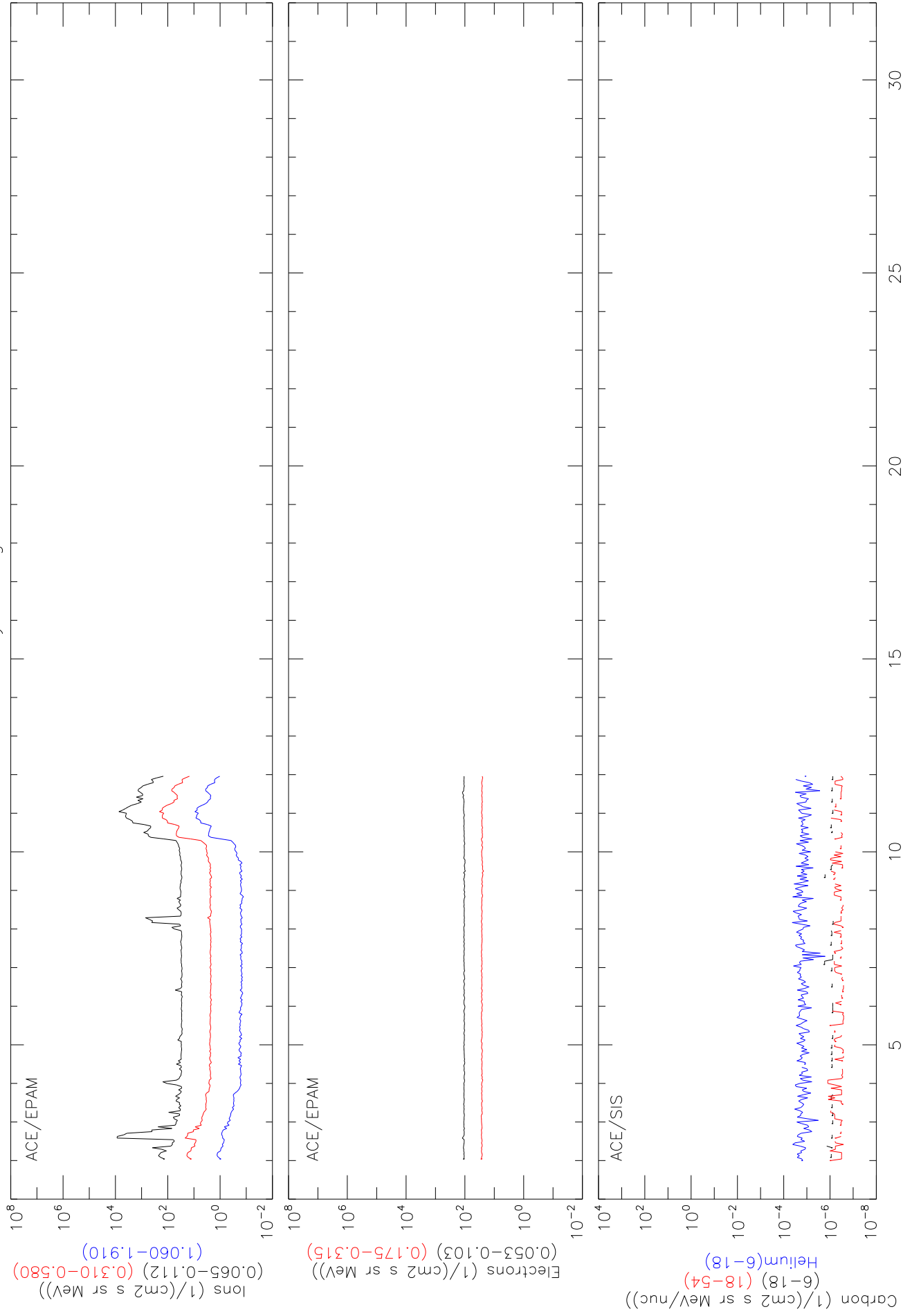
DAYS OF FEBRUARY 2008

ACE LEVEL2 DATA    Solar Wind Plasma    Hourly Averages for FEBRUARY 2008, from SWEPAM



DAYS OF FEBRUARY 2008

Solar Energetic Particles  
ACE LEVEL2 DATA Hourly Averages for FEBRUARY 2008



# SOLAR CORONAL MASS EJECTIONS (CMEs) FROM SOHO/LASCO

<http://cdaw.gsfc.nasa.gov/>

Center for Solar Physics and Space Weather (CSPSW) – The Catholic University of America/NRL/NASA  
FEBRUARY 2008

First C2 Appearance		Central Width			Linear Fit			Measurement		Remarks
Date	Time UT	Position Angle degree	Angular Width degree	Speed km/s	Initial km/s	Final km/s	20R km/s	Accel m/s <sup>2</sup>	Position Angle degree	
2008/02/01	00:30:04	59	19	186	198	173	149	-0.8*	69	Poor Event
2008/02/01	07:31:38	299	12	222	272	171	0	-12.0*	290	Poor Event
2008/02/01	17:54:04	116	34	194	143	247	552	11.8*	116	Very Poor; Only C2
2008/02/02	05:30:04	62	21	135	87	179	538	11.8*	65	Very Poor; Only C2
2008/02/02	07:31:41	134	11	252	242	262	312	1.6*	128	Very Poor Event
2008/02/02	13:31:44	348	4	256	285	227	0	-17.2*	348	Very Poor; Only C2
2008/02/03	05:30:04	121	8	195	108	292	1405	85.0*	117	Very Poor; 3pts; Only C2
2008/02/03	10:34:03	121	9	215	103	326	966	39.0*	116	Poor Event; Only C2
2008/02/04	00:30:04	80	71	364	296	433	511	7.1*	71	Very Poor Event
2008/02/04	02:06:04	310	26	255	231	280	294	1.4*	310	Poor Event
2008/02/04	09:54:04	97	164	306	330	277	275	-1.4*	106	Poor Event; Partial Halo
2008/02/04	22:31:46	23	5	179	139	221	651	17.0*	28	Very Poor; Only C2
2008/02/05	03:06:04	273	11	302	391	208	0	-34.8*	268	Very Poor; Only C2
2008/02/05	05:30:04	257	59	368	316	420	529	7.4*	265	
2008/02/05	06:30:04	138	12	635	451	831	1760	135.2*	131	Poor Event
2008/02/05	10:06:04	147	43	226	152	297	587	13.4*	138	Poor Event; Only C2
2008/02/05	11:18:04	67	9	190	134	244	303	3.5*	74	Very Poor; Only C3
2008/02/05	15:06:05	44	39	127	109	144	227	1.8*	50	Very Poor Event
2008/02/06	12:54:06	71	20	215	147	284	407	6.8*	70	Very Poor Event
2008/02/07	20:58:39	33	7	193	40	354	1496	94.3*	40	Very Poor; 3pts; Only C2
2008/02/07	23:30:04	156	8	98	188	18	0	-32.6*	150	Very Poor; 3pts; Only C2
2008/02/08	02:30:04	229	19	203	119	298	328	4.4*	231	Poor Event
2008/02/08	06:54:05	346	7	314	394	235	0	-43.7*	347	Very Poor; Only C2
2008/02/08	08:54:04	348	12	607	774	459	0	-48.0*	346	Poor Event
2008/02/08	09:54:04	237	33	194	161	229	251	1.7*	233	Poor Event
2008/02/09	02:30:04	224	12	388	353	426	482	4.8*	232	Poor Event
2008/02/09	05:54:04	106	18	285	232	343	615	13.6*	97	Poor Event
2008/02/09	16:30:04	226	6	581	553	609	614	3.1*	232	Poor Event
2008/02/09	19:54:05	144	7	163	161	165	235	1.2*	139	Very Poor; Only C2
2008/02/10	09:54:04	166	4	452	512	397	0	-52.8*	164	Very Poor; 3pts; Only C2
2008/02/10	18:54:04	221	7	324	240	403	488	7.9*	224	Poor Event
2008/02/11	00:54:05	288	5	313	347	278	208	-3.3*	288	Very Poor Event
2008/02/12	01:30:04	243	67	266	68	452	394	6.3*	239	
2008/02/13	10:30:21	54	130	157	90	225	237	2.3*	83	Poor Event; Partial Halo
2008/02/15	01:54:04	90	53	163	0	364	340	5.0*	90	
2008/02/16	08:30:04	242	10	379	557	203	0	-70.7*	246	Very Poor; Only C2
2008/02/16	10:30:20	283	24	284	228	346	390	5.0*	290	Very Poor Event
2008/02/16	16:30:04	77	26	189	107	282	287	3.0*	79	
2008/02/17	00:30:04	80	13	226	204	249	276	1.5*	83	Poor Event
2008/02/17	02:30:04	120	40	130	154	106	0	-7.8*	118	Very Poor; Only C2
2008/02/17	04:30:04	159	6	354	13	728	2841	331.2*	155	Very Poor; 3pts; Only C2
2008/02/17	07:30:05	297	37	203	78	315	414	8.2*	288	Very Poor Event
2008/02/17	12:54:04	356	3	192	110	280	1077	47.3*	357	Very Poor; 3pts; Only C2

# SOLAR CORONAL MASS EJECTIONS (CMEs) FROM SOHO/LASCO

<http://cdaw.gsfc.nasa.gov/>

Center for Solar Physics and Space Weather (CSPSW) – The Catholic University of America/NRL/NASA  
FEBRUARY 2008

First C2 Appearance		Central Width			Linear Fit			----2nd order speed----	Accel	Measurement	
Date	Time UT	Position Angle degree	Angular Width degree	Speed km/s	Initial km/s	Final km/s	20R km/s	m/s <sup>2</sup>	Position Angle degree	Remarks	
2008/02/17	12:54:04	237	5	129	113	146	481	9.1*	238	Very Poor; 3pts; Only C2	
2008/02/17	12:54:04	142	9	174	36	321	1237	65.8*	134	Very Poor; Only C2	
2008/02/18	03:30:04	355	6	324	226	431	1527	94.7*	357	Very Poor; 3pts; Only C2	
2008/02/18	10:54:04	289	13	195	146	245	335	4.4*	287	Very Poor Event	
2008/02/19	00:30:04	233	7	451	506	398	0	-25.2*	239	Poor Event	
2008/02/23	21:24:04	60	64	246	106	398	333	4.8	71		
2008/02/24	10:00:04	206	6	486	593	380	0	-100.2*	210	Very Poor; Only C2	
2008/02/24	10:12:05	74	25	378	221	565	512	9.0*	91		
2008/02/25	08:00:04	125	9	352	402	295	0	-9.2*	125	Poor Event	
2008/02/25	16:12:06	267	16	743	168	1348	1839	137.4*	273	Very Poor Event	
2008/02/25	21:12:07	68	5	309	302	317	334	1.1*	72	Very Poor Event	
2008/02/25	23:12:07	70	8	214	159	269	448	8.4*	71	Very Poor Event	
2008/02/26	02:00:04	64	4	290	240	343	483	8.5*	69	Very Poor Event	
2008/02/26	15:24:04	276	53	126	38	218	204	1.7*	272	Poor Event	
2008/02/27	01:18:04	68	7	196	203	188	168	-0.7*	69	Very Poor Event	
2008/02/28	10:50:04	66	16	223	214	232	252	0.8*	68	Very Poor Event	
2008/02/28	19:48:04	273	54	253	170	348	547	11.0*	270	Very Poor Event	
2008/02/29	01:26:04	270	54	195	133	257	514	10.2*	264	Very Poor Event	
2008/02/29	06:26:04	266	48	125	128	123	112	-0.2*	265	Very Poor Event	
2008/02/29	13:26:04	268	35	187	197	178	0	-2.2*	271	Very Poor; Only C2	
2008/02/29	15:06:04	60	7	306	245	367	412	5.2*	65	Very Poor Event	
2008/02/29	21:54:04	109	5	227	325	132	0	-45.3*	109	Very Poor; Only C2	

\* Acceleration is uncertain due to either poor height measurement or a small number of height-time measurements.

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