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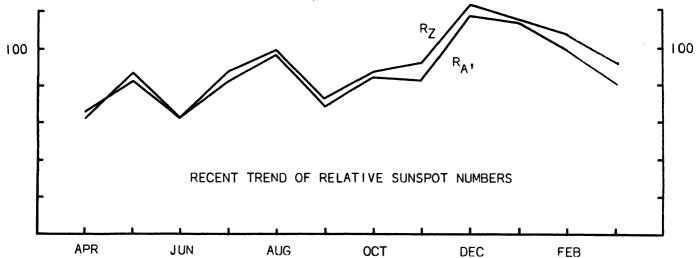
March 1968

## SOLAR ACTIVITY DURING MARCH

Although two ionospheric disturbances were recorded by an AAVSO observer on the 2nd, most of such activity during March occurred at the end of the month from the 21st through the 27th. At this time two flare-producing sunspot groups were developing. One was first seen on the 19th as an A-type group at about 14 degrees south latitude and 60 degrees east of the central meridian. The other started as a B-type group of about 15 very small spots on the 21st near the central meridian at about 11 degrees north latitude. A recording of an ionospheric disturbance associated with one of the flares of the 25th is reproduced on page two.

Sunspot activity continued its downward trend of the previous two months to reach a mean value for March of 80.6 for the American sunspot numbers. The highest numbers occurred toward the end of the month when about 10 groups were present on most days. There were no outstanding large groups during March although there were two spot associations that proved difficult to devide into groups. On the 16th a small bunch of spots formed between two old J-type spots and developed into a small bipolar group. By the 20th the leader of this bipolar group had decayed leaving the follower close enough to the old J spot for the two to resemble a bipolar group which their past history clearly showed was not the case. On 24 March, a large H-type spot came over the northeast limb followed closely by a slightly smaller spot at a somewhat lower latitude. These two resembled a bipolar group but magnetic data revealed that actually it was two spots of the same polarity, both leaders. During disk passage of these two groups, many small spots continually formed and decayed in their vicinity to further complicate their classification.

Development of the above two puzzling groups plus the two groups associated with the ionospheric disturbances is shown schematically from day to day in the March sunspot distribution summary which is sent with this issue of the Solar Bulletin.

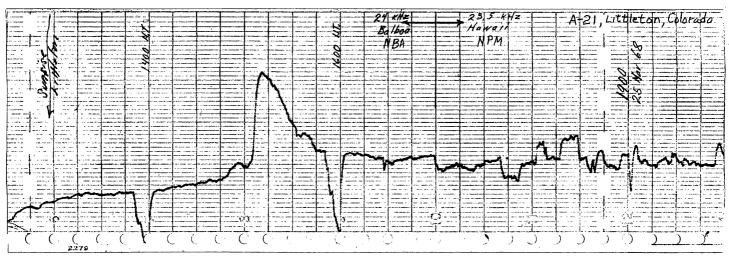


## AMERICAN ( $R_A$ ,) AND ZURICH ( $R_Z$ ) RELATIVE SUNSPOT NUMBERS, MARCH 1968

	**	<b>∠</b>	All for the		
day	$\mathtt{R}_\mathtt{A}$ '	$\mathtt{R}_{\mathbf{Z}}$	day	${f R}_{f A}$ ,	$^{ m R}{ m Z}$
1 2 3 4 5	96 89 75 47 48	108 118 83 92 71	16 17 18 19 20	50 57 46 43 56	46 59 59 53 56
6 7 8 9 10	34 51 55 48 69	56 52 79 69 74	21 22 23 24 25	91 99 100 109 117	70 91 113 132 156
11 12 13 14 15	90 88 72 57 54	85 87 89 67 59	26 27 28 29 30 31	108 128 130 135 135 122	141 146 138 127 154 134
March	mean R <sub>A</sub> ,	= 80.6	March n	nean R <sub>Z</sub> =	92.4

## SUDDEN IONOSPHERIC DISTURBANCES RECORDED DURING MARCH

DAY	MAX.	SEA	SES	DEF.	OBSERVERS	D	ΑY	MAX.	SEA	SES	DEF.	OBSERVERS
2 21 21	1430	3	1 1+ 1+	2 4 5	A-21 A-21 A-21 A-1, 21 A-21	2	5	1509	3	3	5	A-1,21,20 A-21,1,20 A-21,20



A sudden enhancement of signal strength (SES) of very-low-frequency station NBA in Panama caused by a sudden ionospheric disturbance starting at 1505 UT.