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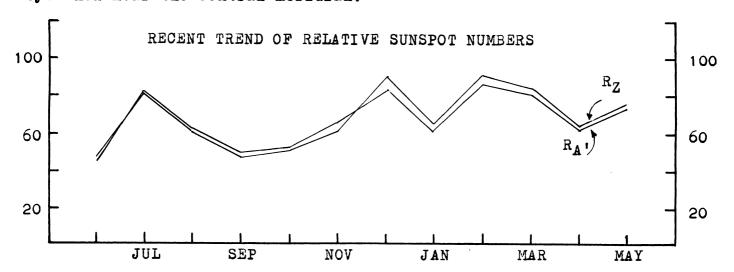
Twenty-eight ionospheric disturbances were recorded by the Solar Division's observers. A large number of high energy events further emphasized this high activity. The event of the 28th possibly equaled the magnitude of any event recorded in recent years.

The event of the 17th as recorded by one of our newer observers, Mr. Allen S. Clarke, Paeonian Springs, Virginia, is shown on page two. This reproduction illustrates that events can be recorded quite well even under adverse conditions of rather high thunderstorm activity and during the ending of the daylight hours when normal attenuation of response occurs.

The event of the 24th reproduced on page two presents a very puzzling point. It was recorded by five SES observers, over ten different signal paths using five different frequencies from 17.8 kHz to 37.2 kHz, with traces giving both high amplitude and a very long decay time, yet none of the seven SEA observers who were in operation at the time appeared to record this disturbance at 27 kHz. It is common for the SES method to record minor events not recorded by the SEA method due to the slightly greater sensitivity of the SES method but for major events, this has not been noted before. Forrelation with other methods should be interesting.

The mean of the American sunspot numbers rose to 75.7 for May with spottedness typical of solar cycle maximum occurring at the middle of the month when at one point thirteen sunspot groups were present on visible disk.

Among interesting groups this month was a high northern latitude group that formed about one day east of the central meridian on the 7th. This developed into an impressive group by west limb passage. It was seen to reappear at the east limb on the 29th. Another impressive group rotated over the east limb on the 24th. Magnetic polarities were badly mixed in this group and it was classified gamma or beta-gamma during its passage across the disk throughout May. It was visible to the unaided eye when near the central meridian.



AMERICAN (R_A ,) AND ZURICH (R_Z) RELATIVE SUNSPOT NUMBERS, MAY 1972

DAY	R _A 1	$^{ m R}{f z}$		DAY	R _A ,	$\mathtt{R}_{\mathbf{Z}}$
1	22	17		16	128	125
2	38	39		17	123	140
3	19	32		18	126	133
4	27	27		19	119	127
5	44	49		20	95	105
6	71	68	$R_{A} = 7$	ans 21	70	89
7	70	79		22	64	77
8	60	73		5.7 23	47	59
9	88	99		24	55	58
10	90	91		8.1 25	59	56
11 12 13 14 15	87 98 127 138 127	79 92 103 130 113		26 27 28 29 30 31	56 40 46 68 80 65	55 41 44 58 79 85

SUDDEN IONOSPHERIC DISTURBANCES RECORDED DURING MAY 1972

DAY	MA X	SEA	SES	DEF	OBSERVERS	DAY	MAX	SEA		DEF	
4	1511		1	5	A1,19,21	15	1951	1+	2	5	A1,8,18,19,21,22,
7	1459	1 -	1	5	A1,19,21,26,31,34						26,30,31,34
7	1845		1 -	5	A19,21,30		0322		2		A31
8	1600		1 -	5	A1,19,21	17	2242	2	2+	5	A1,8,18,19,21,26,
9	1558		1 -	5	A1,19,21						30,31,34
9	1855		1 -	5	A1,19,21	18	1411	2	2-	5	A1,8,18,19,21,22,
11	1637		1 -	5	A1, 19, 21, 30, 31						26,31,32,34
11	1727	1	1	5	A1,19,21,26,30,31	18	1620	1+	1+	5	A1,8,18,19,21,26,
					32						30,31,34
12	1305	1	1	5	A1,8,18,19,22,26	18	2031	1	1+	5	A1,8,18,19,21,26,30
					32,34	18	2237		1 -	5	A19,21,30,31
12	2003		1 -	5	A1,19,21,30	20	1702	1	1+	5	A1,19,21,26,31,34
13	1359	1	1		A1,18,19,21,22,26,	22	2112		1	5	A1,19,21,30,31
					31,34	23	1555	1 -	1	5	A1,19,26,30,31
14	1507		1 -	4	A1,19,21	24	1800		2	5	A1,19,21,30,31
14	2256		1 -	5	A21,30	27	0145		1	4	A31
15	1526		1 -		A19,21	28	1332	3-	3-	5	A1,8,15,18,19,21,
15	1800		1 -	5	A1,21,30,31						26,30,31,32,34

